

Research Article

Mantle wedge enrichment beneath southern Tibet during the late stage (100–45 Ma) of oceanic subduction: Geochemical constraints from mantle-derived intrusions

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ABSTRACT

The Miocene K-rich mafic rocks in the Lhasa Terrane of the southern Tibetan Plateau are considered to be related to either the subduction of the Neo-Tethyan oceanic slab or Indian continental slab. In order to investigate how subduction of the Neo-Tethyan oceanic slab has geochemically affected the mantle beneath southern Tibet, we obtained geochronological and geochemical data for the late Early Cretaceous Cangguo gabbros (97 Ma) and early Eocene Langka diorites (51 Ma) in southern Tibet. Although both the Cangguo gabbros and Langka diorites are characterized by arc-like trace element patterns, with enrichment in light rare earth and large-ion lithophile elements (e.g., Ba, Pb, and U), depletion in high-field-strength elements (e.g., Nb and Ta), and depleted Sr–Nd–Hf isotopic compositions, there are some other major geochemical differences (e.g., Ba/La and Th/La ratios) between these two suites of rocks. These geochemical characteristics of the Cangguo gabbros and Langka diorites indicate they were produced by melting of the depleted mid-ocean ridge mantle (DMM) that was enriched by oceanic sediment-derived fluids and melts, respectively. Our data and previous studies suggest that dehydration and melting of subducted Neo-Tethyan oceanic sediments beneath the Lhasa Terrane occurred in the late stage (100–45 Ma) of Neo-Tethyan oceanic slab subduction, which may have enriched the mantle beneath the Lhasa Terrane. However, the mantle beneath the Lhasa Terrane that was enriched by the subduction of oceanic slab during 100–45 Ma is still isotopically more depleted than the source of Miocene post-collisional K-rich mafic rocks in the Lhasa Terrane. This indicates that the enriched component in the mantle source of the K-rich mafic rocks cannot be simply attributed to materials derived from the Neo-Tethyan oceanic slab, and other additional materials are required.

1. Introduction

Continent–continent collisional orogenic belts are typically associated with the subduction of oceanic lithosphere, closure of oceanic basins, continent–continent collision, and subsequent subduction of continental lithosphere (e.g., Castro et al., 2013; Ernst, 2005). In a typical continent–continent collisional belt, both oceanic and continental crustal materials are recycled into and interact with the overlying mantle (e.g., Zheng and Chen, 2016). In this scenario, the mantle wedge beneath a continent–continent collisional orogen can record crust–mantle interactions in both the continental and earlier oceanic subduction channels (e.g., Zheng and Chen, 2016). The southern Tibetan Plateau is a typical continent–continent collisional orogeny, and records

subduction of Neo-Tethyan oceanic and Indian continental lithosphere (e.g., Chung et al., 2005). The Miocene (24–8 Ma) post-collisional K-rich (ultrapotassic–potassic) mafic rocks are widely distributed in the Lhasa Terrane and are unique mantle-derived magmas produced after the collision between India and Asia. These K-rich mafic rocks have arc-like trace element patterns and highly enriched Sr – Nd isotopic compositions (e.g., Turner et al., 1996; Miller et al., 1999; Ding et al., 2003; Gao et al., 2007; Guo et al., 2013; Liu et al., 2015). However, because the mantle wedge beneath the Lhasa Terrane might have been affected by two different episodes of subduction (i.e., the northward subduction of the Neo-Tethyan oceanic slab and subsequent indentation of the Indian continental lithosphere), it is still unclear how this isotopically enriched mantle source formed. For example, some studies have proposed that

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this isotopically enriched mantle formed by the subduction of Indian continental crust (e.g., Ding et al., 2003; Guo et al., 2013), whereas other studies have proposed it resulted from introduction of oceanic slab-derived materials during Neo-Tethyan subduction (Gao et al., 2007; Liu et al., 2015; Tommasini et al., 2010). Thus, understanding the relationship between the subduction of the Neo-Tethyan oceanic slab and its geochemical effects on the mantle beneath the Lhasa Terrane is the key to resolving this controversy. In general, a subducted oceanic slab consists of oceanic crust and overlying sediments, and there are trace element and isotopic differences between these two components (Chauvel et al., 2009; Plank, 2014; Vervoort et al., 1999, 2011). Previous studies have demonstrated that both oceanic crust and overlying sediments can derive aqueous fluids and/or melts that can enrich the mantle wedge (Carpentier et al., 2009; Labanieh et al., 2010). However, the nature of enrichment in a particular subduction system depends mainly on its thermal structure (Watt et al., 2013). For example, slab-derived fluid-mediated enrichment has prevailed in relatively cold subduction zones, such as the Izu–Bonin–Mariana arcs (e.g., Hauff et al., 2003). In contrast, sediment-derived, melt-mediated enrichment is dominant in relatively hot subduction zones, including the Lesser Antilles and Banda arcs (e.g., Labanieh et al., 2010; Nebel et al., 2011).

It has previously been proposed that initial northward subduction of the Neo-Tethyan oceanic slab might have occurred in the Late Triassic–Early Jurassic (210–174 Ma; Zhu et al., 2008, 2011; Ji et al., 2009; Kang et al., 2014). Subsequently, the initial collision of India and Asia occurred during 65–55 Ma (e.g., Hu et al., 2016; Zhu et al., 2015). During the initial collision of India and Asia, the oceanic slab continued to subduct and dehydrate, thereby producing fluids and melts, which resulted in normal continental arc magmatism and enrichment of the overlying mantle (e.g., Huang et al., 2019, 2020; Zhu et al., 2015). Previous studies have argued that the termination of oceanic slab subduction finally occurs when the slab pull force disappears (e.g., slab break-off; Davies and von Blanckenburg, 1995). Hence, slab break-off of Neo-Tethyan oceanic crust might record the termination of oceanic

subduction in the mantle beneath southern Tibet, which occurred during 53–45 Ma (e.g., Ji et al., 2009, 2016; Leech et al., 2005; Zhu et al., 2013, 2015). Therefore, Mesozoic–Cenozoic (210–45 Ma) mantle-derived rocks in the Lhasa Terrane are suitable for investigating the geochemical effects from Neo-Tethyan oceanic slab subduction on the mantle beneath the Lhasa Terrane. Most studies have focused on Late Triassic–Early Jurassic mafic rocks (e.g., Late Triassic Quxu gabbros, Jurassic Yeba Formation, and Sangri Group) in the Lhasa Terrane, and concluded that oceanic crust-derived, fluid-mediated enrichment was predominant at this time. As such, these mafic rocks tend to have nearly mid-ocean ridge basalt (MORB)-like depleted Sr – Nd – Hf isotopic compositions, and high Ba/La and low Th/Nd ratios (e.g., Kang et al., 2014; Meng et al., 2015; Zhu et al., 2008). However, it is unclear how the mantle beneath the Lhasa Terrane was geochemically modified during the Late Cretaceous–early Eocene (100–45 Ma) by Neo-Tethyan oceanic slab subduction, mainly due to the rarity of exposed mantle-derived rocks during this time. This study presents the detailed petrological, geochronological, major and trace element, and Sr–Nd–Hf isotopic data for two mantle-derived intrusions generated in the late stage (100–45 Ma) of Neo-Tethyan oceanic slab subduction in the Lhasa Terrane. We use our data, along with those from previous studies, to: (1) constrain the source and petrogenesis of these rocks; (2) assess how the mantle source was geochemically modified by subduction of the oceanic slab; (3) explore how the enriched source of the post-collisional K-rich mafic rocks formed.

2. Geology of the study area and sampling

The Tibetan Plateau consists of the Songpan–Gangzi, Qiangtang, Lhasa, and Himalaya Terranes from north to south. The Lhasa Terrane is bounded by the Bangong–Nujiang suture zone (BNSZ) to the north and the Indus–Yarlung Zangbo suture zone (IYZS) to the south (Fig. 1a). It is generally accepted that BNSZ marks the closure of the Bangong–Nujiang Ocean during the Late Cretaceous to Early Cretaceous,

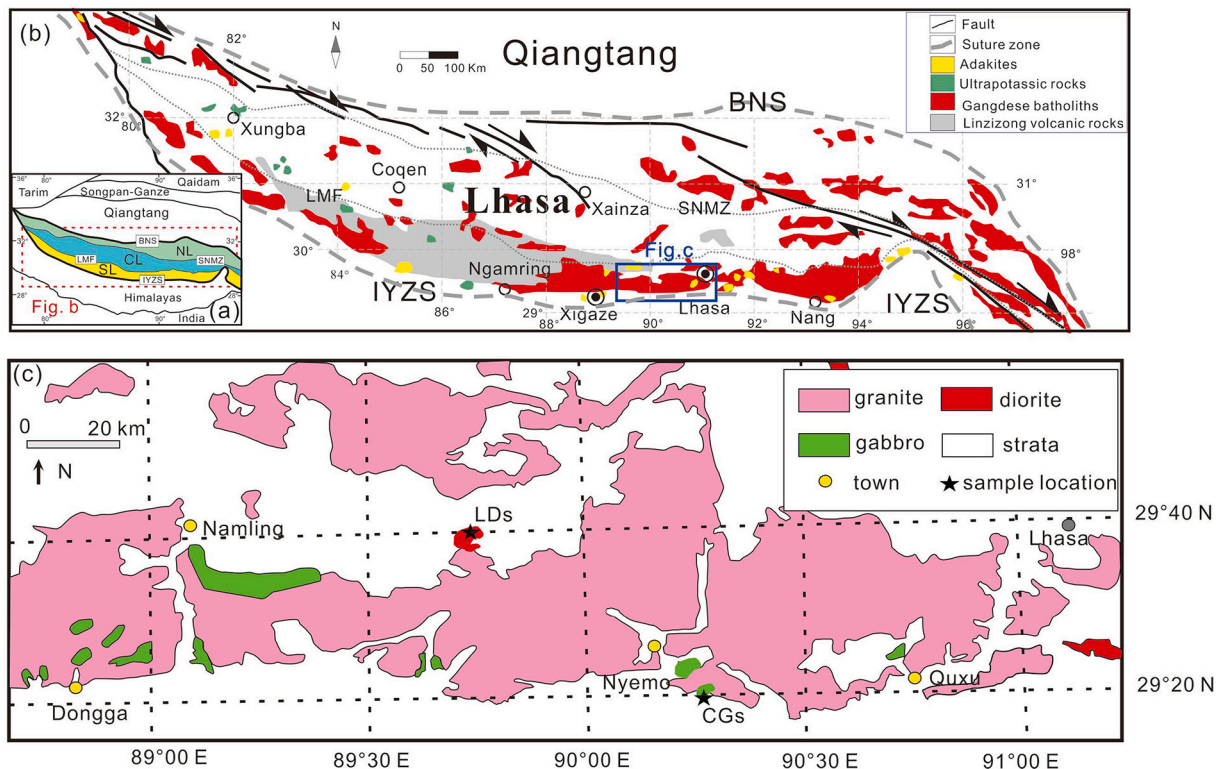


Fig. 1. (a) The Lhasa Terrane in the context of the Tibetan Plateau. (b) Geological map of the Lhasa Terrane, modified from Huang et al. (2019). (c) Simplified map of sample locations in the Lhasa Terrane. Abbreviations: BNS = Bangong–Nujiang Suture; IYZS = Indus–Yarlung Zangbo Suture.

whereas the IYZSZ marks the closure of the Neo-Tethyan Ocean during the Late Cretaceous to Early Paleogene (Dewey et al., 1988; Yin and Harrison, 2000). The Lhasa Terrane could be subdivided into the northern, central, and southern subterrane separated by the Shiquan River–Nam Tso Mélange zone (SNMZ) and Luobadui–Milashan fault (LMF), respectively (Fig. 1b, Zhu et al., 2011).

The northern Lhasa subterrane is dominated by juvenile crust covered by the Middle Triassic to Cretaceous sedimentary rocks along with Early Cretaceous volcanic rocks (andesite, dacite, rhyolite, and volcanoclastic rocks) and associated granitoids (Zhu et al., 2011, 2013). The central Lhasa subterrane is commonly considered as an ancient microcontinent with Precambrian basement rocks (Zhu et al., 2011, 2013). The Precambrian basement was covered with Ordovician, Silurian, Permian–Carboniferous, Triassic, and Late Jurassic–Early Cretaceous volcano-sedimentary rocks (Zhu et al., 2011, 2013). The early Cretaceous magmatic rocks in this subterrane were extending ~1200 km and yielded ages of 140–100 Ma (Zhu et al., 2011). The southern Lhasa subterrane may represent the southernmost part of the Asian lithosphere. The southern Lhasa subterrane is considered to be the juvenile continental crust dominated by the Late Triassic to Eocene (205–40 Ma) Gangdese batholith (Chung et al., 2005; Ji et al., 2009), the Mesozoic Yeba Formation, the Sangri Group, the Paleogene Linzizong volcanic succession, and post-collisional (ultra)potassic and adakitic rocks (Chung et al., 2005; Ji et al., 2009; Mo et al., 2007, 2008). Both the Cangguo gabbros and Langka diorites were sampled in the central part of southern Lhasa subterrane (Fig. 1b).

The Cangguo gabbros (CGs) are located in the Cangguo county, 10 km southeast of Nyemo town (Fig. 1c). The gabbros are dark gray, fine- to medium-grained, and consist of hornblende, plagioclase, clinopyroxene with accessory zircon and apatite (Fig. 2a–b).

The Langka diorites (LDs) lie in the Langka county, 70 km northwest

of Nyemo town (Fig. 1c). These rocks are medium- to coarse-grained and mainly consist of hornblende, plagioclase, and biotite along with accessory Fe–Ti oxides (Fig. 2c–d).

3. Analytical techniques

3.1. Zircon U–Pb dating

Zircons were extracted from fresh samples following normal separation methods, and they were then handpicked and mounted together onto double-sided adhesive tape before being enclosed in epoxy resin discs. Zircon U–Pb analyses of samples 19CG–08 and 19LK–07 were conducted in an Agilent 7700e ICP–MS system equipped with a 193 nm laser, at Sample Solution Analytical Technology Co., Ltd. (Wuhan, China). The zircon 91,500 was used as the external standard, and NIST SRM 610 were used as the reference material. The detailed analytical technique has been outlined by Zong et al. (2017). The time-resolved spectra were analyzed off-line by ICPMSDataCal software. The common-Pb correction followed the method described by Andersen (2002). The Isoplot/Ex 3.0 software (Ludwig, 2003) was used for calculating U–Pb ages.

3.2. Whole-rock major and trace element analyses

All samples were crushed to select fresh chips. The whole-rock major, trace element and Sr–Nd–Hf isotope analyses were also determined at Sample Solution Analytical Technology Co., Ltd. (Wuhan, China). The fresh chips were powdered in an agate mill to 200 mesh. Whole-rock major and trace element analyses were determined by X-ray fluorescence (XRF) spectrometry and inductively coupled plasma mass spectrometry (ICP–MS), respectively. For major elements, the analytical

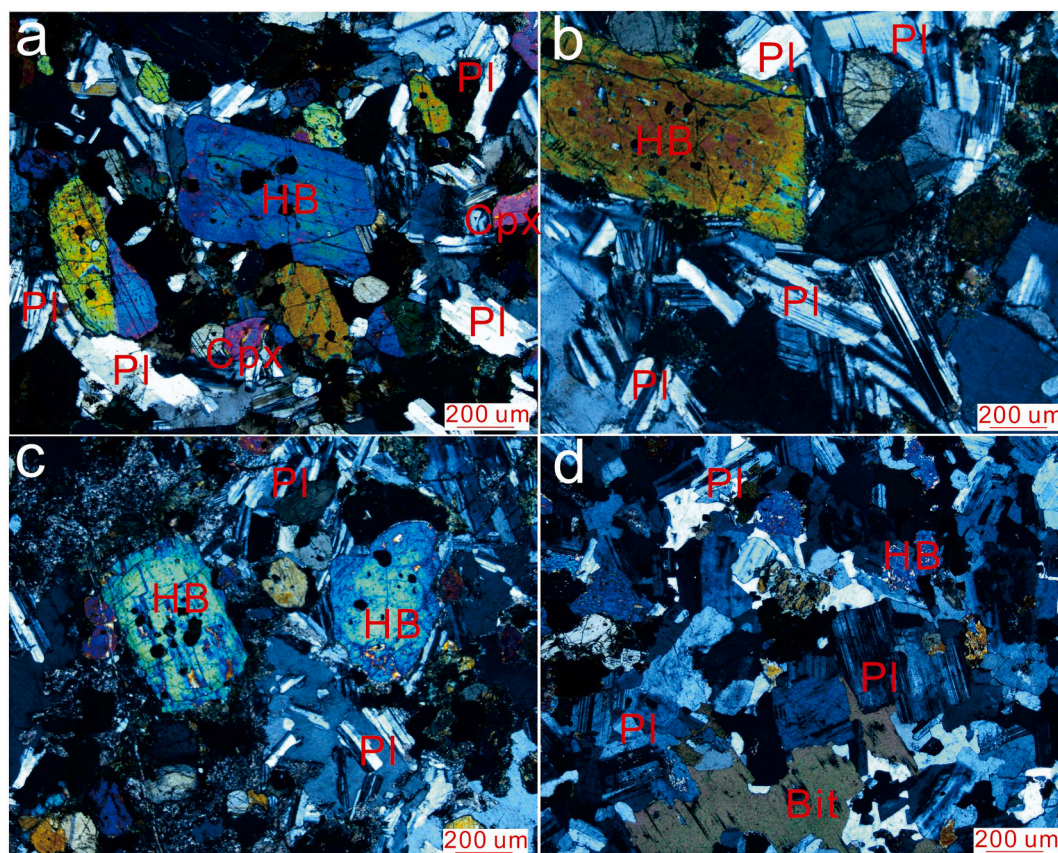


Fig. 2. Photomicrographs of the CGs and LDs. (a) and (b) clino- and hornblende phenocrysts from the CGs. (c) and (d) hornblende, plagioclase, and biotite phenocrysts from the LDs. Abbreviations here: Cpx, clinopyroxene; Pl, plagioclase; Hb, hornblende; Bit, biotite.

errors are within 5%. Precision for trace elements is better than 10%, based on repeated analysis of USGS standards AGV-2, BCR-2 and BHVO-2. The measured values for rock standards are provided in Supplementary table 1.

3.3. Sr–Nd–Hf isotope analyses

The Sr–Nd–Hf isotopes were measured using a Thermo-Fisher Neptune multicollector ICP-MS system. For the Sr–Nd isotopic analysis, whole powders were dissolved in HF–HClO₄ and the dissolution was conducted at 200 °C for one week. Sr and Nd were separated using conventional ion exchange columns, and the Nd fractions were further separated with HDEHP columns. The measured ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd ratios were normalized to ⁸⁶Sr/⁸⁸Sr = 0.1194 and ¹⁴⁶Nd/¹⁴⁴Nd = 0.7219, respectively. The reported ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd ratios were adjusted to the NBS SRM 987 standard

(⁸⁷Sr/⁸⁶Sr = 0.710241 ± 12, 2σ, n = 8) and the JNdi-1 standard (¹⁴³Nd/¹⁴⁴Nd = 0.512115 ± 7, 2σ, n = 8), respectively. The procedural blanks were 200 pg for Sr and ~ 30 pg for Nd. For the Hf isotopic analysis, about 200 mg of powders were digested in Teflon bombs with a mixture of concentrated HNO₃ + HF and dried on a hot plate. This was followed by the addition of concentrated HNO₃, HF and HClO₄ sealed in bombs and kept in an oven at 190 °C. After cooling, the Teflon bomb was opened and then 1 ml HNO₃ was added and evaporated to dryness. A modified ion exchange single-column of LN-Spec resin was used for Hf separation. Mass discrimination correction was carried out via internal normalization to ¹⁷⁹Hf/¹⁷⁷Hf ratio of 0.7325. During our analysis, replicate measurements of the Hf standard JMC 475 gave ¹⁷⁶Hf/¹⁷⁷Hf of 0.282152 ± 0.000010 (2σ, n = 3). In-house Alfa Hf standard measured during analysis gave an average ¹⁷⁶Hf/¹⁷⁷Hf of 0.282224 ± 0.000015 (2σ, n = 10). The procedural blank was 50 pg for Hf. Isotope ratios of geological reference materials (BCR-2, RGM) measured during the

Table 1
Whole-rock major and trace elements of Cangguo gabbros (CGs) and Langka diorites (LDs).

Sample	19CG-01	19CG-02	19CG-03	19CG-04	19CG-05	19CG-06	19CG-07	19CG-08	19LK-01	19LK-02	19LK-03	19LK-03a	19LK-04	19LK-05	19LK-06	19LK-07
(wt%)																
SiO ₂	49.32	48.91	49.30	49.41	49.23	48.07	49.55	49.00	60.75	60.79	58.83	58.90	61.10	61.04	60.51	58.59
TiO ₂	0.88	0.88	0.88	0.83	0.85	0.90	0.84	0.87	0.79	0.78	0.81	0.80	0.79	0.77	0.78	0.83
Al ₂ O ₃	17.26	16.77	16.49	17.00	16.90	16.49	17.01	16.61	16.01	16.00	16.54	16.48	16.28	16.11	16.10	16.56
TFe ₂ O ₃	10.69	10.70	10.50	10.24	10.37	10.96	10.11	10.40	6.71	6.64	7.27	7.29	6.48	6.43	6.70	7.41
MnO	0.19	0.20	0.19	0.19	0.19	0.20	0.19	0.20	0.11	0.11	0.13	0.13	0.11	0.11	0.11	0.13
MgO	5.05	5.02	5.23	5.00	5.13	5.48	4.89	5.34	2.96	2.91	3.22	3.21	2.71	2.80	2.97	3.24
CaO	10.67	10.69	10.73	10.43	10.47	11.26	10.13	10.88	5.28	5.26	5.95	5.98	5.47	5.24	5.43	5.99
Na ₂ O	2.72	2.71	2.59	2.76	2.93	2.56	2.70	2.90	3.43	3.47	3.67	3.70	3.67	3.50	3.49	3.66
K ₂ O	2.43	2.53	2.56	2.78	2.39	2.01	2.96	1.99	3.09	3.07	2.36	2.40	2.49	3.06	2.90	2.42
P ₂ O ₅	0.42	0.48	0.45	0.47	0.46	0.53	0.43	0.45	0.20	0.19	0.22	0.21	0.21	0.19	0.20	0.21
Mg#	52.4	52.2	53.7	53.2	53.5	53.8	53.0	54.5	50.7	50.6	50.8	50.7	49.4	50.4	50.8	50.5
LOI	0.08	0.42	0.65	0.30	0.52	0.91	0.56	0.81	0.44	0.33	0.62	0.55	0.53	0.40	0.40	0.49
total	99.69	99.32	99.57	99.37	99.45	99.38	99.35	99.44	99.75	99.54	99.61	99.64	99.84	99.64	99.59	99.51
Li	6.66	13.7	9.57	9.10	14.6	16.8	12.3	8.57	20.4	18.1	19.2	20.0	18.2	19.2	19.2	19.2
Be	0.90	1.15	1.07	1.50	1.26	1.05	1.25	1.11	1.45	1.54	1.48	1.78	1.59	1.37	1.64	1.64
Sc	24.2	25.1	25.8	24.7	24.8	27.5	23.5	26.2	15.5	15.6	17.8	14.9	14.7	15.6	18.0	18.0
V	314	319	318	308	318	334	302	315	142	144	159	137	136	142	161	161
Cr	19.3	17.6	28.7	18.9	18.6	23.4	20.2	27.8	27.5	28.8	30.5	22.2	25.3	29.2	31.8	31.8
Co	34.5	35.5	36.5	33.8	37.0	38.8	32.5	34.9	19.6	19.8	21.6	18.8	18.5	19.8	21.8	21.8
Ni	22.2	22.9	22.5	21.3	23.1	23.9	21.4	23.9	17.4	17.9	19.6	14.9	16.2	17.2	19.5	19.5
Cu	128	169	146	142	117	167	87.6	89.9	44.5	49.5	47.3	52.8	44.6	48.4	41.5	41.5
Zn	83.0	89.6	84.6	84.5	85.4	91.7	83.9	83.6	74.9	74.2	77.1	73.4	68.1	70.1	78.2	78.2
Ga	17.8	18.4	17.4	18.1	18.2	18.1	17.9	18.1	17.8	18.0	18.7	19.2	17.7	17.8	18.2	18.2
Rb	36.3	44.1	39.6	48.0	53.2	30.0	51.0	34.2	107	107	84.2	96.5	103	97.8	83.7	83.7
Sr	1268	1231	1185	1205	1224	1178	1250	1154	472	481	526	473	473	482	526	526
Y	19.9	23.5	21.7	22.5	21.2	23.4	21.0	22.2	23.2	23.8	24.0	25.7	22.9	23.1	24.3	24.3
Zr	39.9	55.8	39.6	55.6	37.6	48.4	57.7	41.8	177	178	144	187	176	154	137	137
Nb	1.49	2.46	1.26	2.43	1.78	1.48	2.35	1.53	8.52	9.10	7.89	9.35	8.21	8.19	9.02	9.02
Sn	0.66	0.75	0.70	0.73	0.73	0.76	0.72	0.76	1.38	1.36	1.34	1.53	1.32	1.28	1.30	1.30
Cs	0.80	2.39	1.00	2.36	1.24	1.36	1.90	0.66	8.28	7.91	5.85	5.81	7.82	7.28	6.32	6.32
Ba	1153	1194	1316	1276	1181	984	1390	1020	602	606	548	450	596	595	561	561
La	20.4	23.7	21.6	23.1	22.2	24.0	21.9	21.1	22.9	23.5	23.2	26.3	22.8	22.6	22.7	22.7
Ce	40.1	48.3	43.1	46.3	44.2	48.8	44.3	43.2	47.5	48.8	48.2	55.5	47.1	46.9	47.7	47.7
Pr	5.01	5.97	5.47	5.74	5.41	6.15	5.46	5.52	5.76	5.88	5.87	6.71	5.69	5.71	5.95	5.95
Nd	21.6	25.2	23.8	24.1	22.8	26.5	23.5	23.5	22.7	22.8	23.0	26.0	22.2	22.1	23.5	23.5
Sm	5.23	6.01	5.65	5.83	5.62	6.30	5.71	5.92	4.92	5.10	5.07	5.38	4.73	4.38	4.92	4.92
Eu	1.90	1.93	1.82	2.05	1.87	1.90	1.87	1.92	1.11	1.09	1.22	1.14	1.07	1.10	1.23	1.23
Gd	4.69	5.55	5.02	5.16	4.84	5.69	4.82	5.30	4.00	4.13	4.38	4.54	4.00	3.97	4.49	4.49
Tb	0.67	0.82	0.75	0.70	0.70	0.85	0.74	0.73	0.65	0.68	0.69	0.73	0.61	0.67	0.73	0.73
Dy	3.71	4.06	4.00	3.94	3.76	4.30	3.88	4.10	3.88	3.97	4.02	4.31	3.82	3.95	4.29	4.29
Ho	0.70	0.80	0.79	0.80	0.71	0.84	0.73	0.75	0.81	0.81	0.82	0.88	0.78	0.77	0.85	0.85
Er	1.86	2.16	1.99	2.07	1.99	2.24	1.99	2.07	2.29	2.31	2.35	2.53	2.23	2.21	2.43	2.43
Tm	0.24	0.29	0.27	0.28	0.26	0.29	0.27	0.27	0.32	0.34	0.33	0.37	0.33	0.33	0.35	0.35
Yb	1.59	1.82	1.71	1.73	1.65	1.83	1.65	1.79	2.23	2.24	2.27	2.46	2.16	2.13	2.37	2.37
Lu	0.23	0.27	0.24	0.25	0.24	0.27	0.25	0.25	0.33	0.34	0.34	0.37	0.33	0.32	0.35	0.35
Hf	1.29	1.72	1.41	1.55	1.18	1.52	1.60	1.38	4.68	4.81	4.00	5.29	4.61	4.31	3.97	3.97
Ta	0.082	0.12	0.084	0.14	0.11	0.091	0.12	0.082	0.61	0.65	0.59	0.70	0.64	0.59	0.66	0.66
Tl	0.033	0.058	0.064	0.045	0.13	0.036	0.072	0.058	0.51	0.48	0.36	0.42	0.43	0.40	0.35	0.35
Pb	6.43	7.09	6.44	8.42	6.47	7.03	7.45	6.26	17.9	18.3	17.5	16.4	18.6	18.0	17.8	17.8
Th	1.04	2.28	0.95	2.69	1.47	1.13	2.58	1.19	12.5	10.5	13.4	15.6	12.9	8.44	11.0	11.0
U	0.29	0.64	0.41	0.78	0.53	0.31	0.73	0.56	3.00	3.63	2.76	3.58	3.87	3.33	2.91	2.91

Note: Total iron as TFe₂O₃; Mg# = 100 × Mg²⁺ / (Mg²⁺ + Fe²⁺), assuming Fe³⁺ / (Fe²⁺ + Fe³⁺) = 0.15. LOI = loss on ignition.

analytical procedure are presented in Supplementary table 2.

4. Results

The LA-ICPMS zircon U-Pb geochronology data for the CGs and LDs are given in Supplementary table 3. Whole rock major- trace element and Sr–Nd–Hf isotope data for the CGs and LDs are listed in Table 1 and Table 2, respectively.

4.1. Zircon U-Pb ages

Representative CL images of analyzed zircon and U-Pb concordia diagrams of two samples are shown in Fig. 3. The zircon grains from CGs and LDs are euhedral to subhedral, and show oscillatory zoning in CL images, and have high Th/U values (typically >0.1; Fig. 3; Supplementary table 3), indicating an igneous origin. Therefore, the U-Pb zircon ages represent the timing of crystallization of these magmatic rocks. Zircon grains from one sample of CGs (19CG–08) yield weighted-mean $^{206}\text{Pb}/^{238}\text{U}$ ages of 97.1 ± 0.2 Ma (MSWD = 0.49) (Fig. 3a). Zircon grains from one sample of LDs (19LK–07) yield weighted-mean $^{206}\text{Pb}/^{238}\text{U}$ ages of 50.9 ± 0.4 Ma (MSWD = 1.02) (Fig. 3b).

4.2. Whole-rock major- trace elements and Sr–Nd–Hf isotopes

The CGs have relatively homogeneous MgO (4.89–5.48 wt%), $\text{Mg}^\#$ (52.2–54.5), Ni (21.3–23.9 ppm), K_2O (1.99–2.96 wt%), and SiO_2 (48.1–49.6 wt%) values with high $\text{K}_2\text{O}/\text{Na}_2\text{O}$ ratios (0.69–1.1), belonging to the shoshonite series (Fig. 4). The CGs show enrichment of light rare earth elements (LREEs), large ion lithophile elements (LILEs; e. g., Ba, U, Pb, and Sr) and depletion in high field strength elements (HFSEs; e. g., Nb, Ta, Zr, and Hf) on the chondrite-normalized REE and primitive mantle-normalized multi-elements variation diagrams (Fig. 5). Compared with CGs, the LDs have obviously lower MgO (2.71–3.24 wt%), $\text{Mg}^\#$ (49.4–50.8), Ni (14.9–19.6 ppm), higher SiO_2 (58.6–61.1 wt%), and K_2O (2.36–3.09 wt%) values with slightly lower $\text{K}_2\text{O}/\text{Na}_2\text{O}$ (0.65–0.90) ratios, which belong to the high-K calc-alkaline series (Figs. 4). The LDs show enrichment of LREEs with high (La/Yb)_N values (6.9–7.7) (Fig. 5a) and are characterized by enrichment of Pb, Th, U, and depletion of Nb and Ti on primitive mantle-normalized multi-elements variation diagram (Fig. 5b).

The CGs are characterized by the lower initial $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$ ratios (0.703754–0.703796), higher $\epsilon\text{Nd}_{(t)}$ (+3.5 to +3.8), and $\epsilon\text{Hf}_{(t)}$ values (+11.5 to +12.5) than those of the LDs ($^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$ = 0.704295–0.704340; $\epsilon\text{Nd}_{(t)}$ = +2.8 – +3.0; $\epsilon\text{Hf}_{(t)}$ = +9.3 – +10.1) (Fig. 6). Both of them have positive $\Delta\epsilon\text{Hf}_{(t)}$ (+3.3 – +4.5 for CGs and +2.2 – +3.1 for LDs, $\Delta\epsilon\text{Hf}_{(t)} = \epsilon\text{Hf}_{(t)} - 1.36\epsilon\text{Nd}_{(t)} - 2.89$, Vervoort et al., 1999) and plot above the terrestrial array in an Nd–Hf isotopic space (Fig. 6b).

Table 2

Whole- rock Sr–Nd–Hf isotope for Cangguo gabbros (CGs) and Langka diorites (LDs).

Sample	19CG-01	19CG-02	19CG-03	19CG-04	19CG-05	19LK-01	19LK-02	19LK-03	19LK-04	19LK-05
$^{87}\text{Sr}/^{86}\text{Sr}$	0.703878	0.703905	0.703925	0.703907	0.703946	0.704770	0.704753	0.704638	0.704753	0.704736
2 σ	0.000010	0.000008	0.000011	0.000010	0.000010	0.000013	0.000009	0.000009	0.000008	0.000012
$^{143}\text{Nd}/^{144}\text{Nd}$	0.512801	0.512796	0.512799	0.512795	0.512791	0.512768	0.512771	0.512762	0.512757	0.512770
2 σ	0.000004	0.000004	0.000003	0.000004	0.000004	0.000003	0.000004	0.000006	0.000004	0.000004
$^{176}\text{Hf}/^{177}\text{Hf}$	0.283106	0.283104	0.283088	0.283080	0.283086	0.283034	0.283022	0.283014	0.283025	0.283012
2 σ	0.000007	0.000007	0.000008	0.000007	0.000006	0.000006	0.000006	0.000008	0.000009	0.000008
$^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$	0.703768	0.703767	0.703796	0.703754	0.703779	0.704312	0.704304	0.704314	0.704340	0.704295
$\epsilon\text{Nd}_{(t)}$	3.8	3.7	3.8	3.7	3.5	2.9	3.0	2.8	2.8	3.0
$\epsilon\text{Hf}_{(t)}$	12.4	12.5	11.8	11.6	11.5	10.1	9.6	9.3	9.7	9.3
$\Delta\epsilon\text{Hf}_{(t)}$	4.3	4.5	3.7	3.7	3.7	3.1	2.6	2.5	3.0	2.2

Note: The initial Sr–Nd–Hf isotopic values of studied rocks were calculated using their eruption ages. The present chondrite has: $^{87}\text{Rb}/^{86}\text{Sr} = 0.0847$; $^{87}\text{Sr}/^{86}\text{Sr} = 0.7045$; $^{147}\text{Sm}/^{144}\text{Nd} = 0.1967$, $^{143}\text{Nd}/^{144}\text{Nd} = 0.512638$, $^{176}\text{Lu}/^{177}\text{Hf} = 0.0332$, and $^{176}\text{Hf}/^{177}\text{Hf} = 0.282772$. $\Delta\epsilon\text{Hf}_{(t)} = \epsilon\text{Hf}_{(t)} - 1.36\epsilon\text{Nd}_{(t)} - 2.89$ (Vervoort et al., 1999).

5. Discussion

5.1. Effects of crustal contamination and fractional crystallization

To constrain the nature of the mantle source of the studied rocks, we first assessed the possible effects, if any, due to low-pressure processes such as crustal contamination and fractional crystallization (FC). Based on the following lines of evidence, we suggest that crustal contamination was insignificant. Firstly, no zircon xenocrysts were found in the studied rocks. Secondly, both the CGs and LDs exhibit limited variations in $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$ ratios and $\epsilon\text{Nd}_{(t)}$ values with decreasing MgO contents (Fig. 7a–b). Thirdly, given that CGs and LDs were sampled in the southern Lhasa subterrane which was proven to be the juvenile continental crust with depleted Sr–Nd isotopic values and low Nb/La ratios (e. g., Rudnick and Gao, 2003; Zhu et al., 2013), mantle-derived magmas contaminated by such crustal materials will exhibit decreasing Nb/La with decreasing $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$ and increasing $\epsilon\text{Nd}_{(t)}$. However, $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$ ratios and $\epsilon\text{Nd}_{(t)}$ values of the studied samples do not correlate with Nb/La ratios (Fig. 7c–d), thereby ruling out significant crustal contamination.

Compared with primitive mantle-derived magmas, which have high Ni contents (400 ppm) and $\text{Mg}^\#$ (73–81) values (Wilson, 1989), the low $\text{Mg}^\#$ values (52.2–54.5 for the CGs; 49.4–50.8 for the LDs) and Ni concentrations (21.3–23.9 ppm for the CGs; 14.9–19.6 ppm for the LDs) of the CGs and LDs suggest their parental magmas have undergone some FC. The positive correlations between Ni, CaO, and MgO contents suggest that the CGs experienced FC of olivine and clinopyroxene (Fig. 8a–b). The lack of positive correlations between Al_2O_3 , Sr, and MgO contents indicate that plagioclase fractionation was insignificant (Fig. 8c–d). However, fractionation of plagioclase and hornblende appears to have dominated the evolution of the LDs magmas, as evidenced by the positive correlations between Al_2O_3 , Sr, TFe_2O_3 , Sc, and MgO contents (Fig. 8c–f). This is consistent with the presence of plagioclase and hornblende phenocrysts in the LDs.

Given that crustal contamination was negligible during the generation of the CGs and LDs, the geochemical variations in the two intrusions were largely due to crystal fractionation. Although FC of some minerals (olivine, pyroxene, plagioclase, and hornblende) would cause variations in whole-rock major element (e. g., MgO, Al_2O_3 , and CaO) and trace element (e. g., Sr) contents, this process would not significantly affect the ratios of some incompatible trace elements, such as Ba/La, Th/Sm, and Th/La (e. g., Singer et al., 2007). In the following section, these trace element ratios and Sr–Nd–Hf isotope data for the CGs and LDs are used to constrain the nature of their mantle source.

5.2. Mantle source

5.2.1. Mantle source of the CGs

The CGs were produced in the Gangdese continental arc setting, and

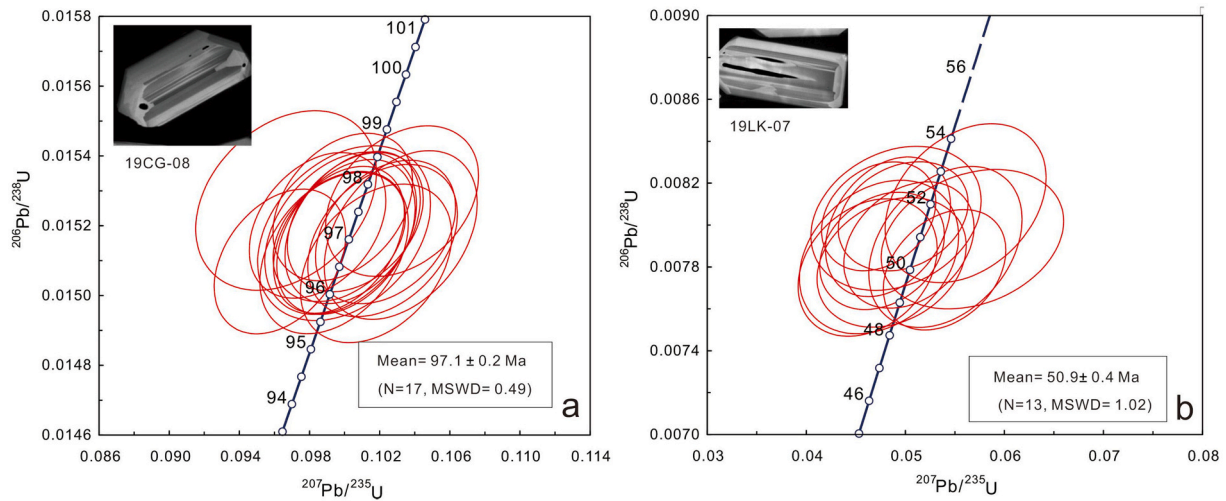


Fig. 3. Zircon U–Pb Concordia diagrams of the CGs and LDs.

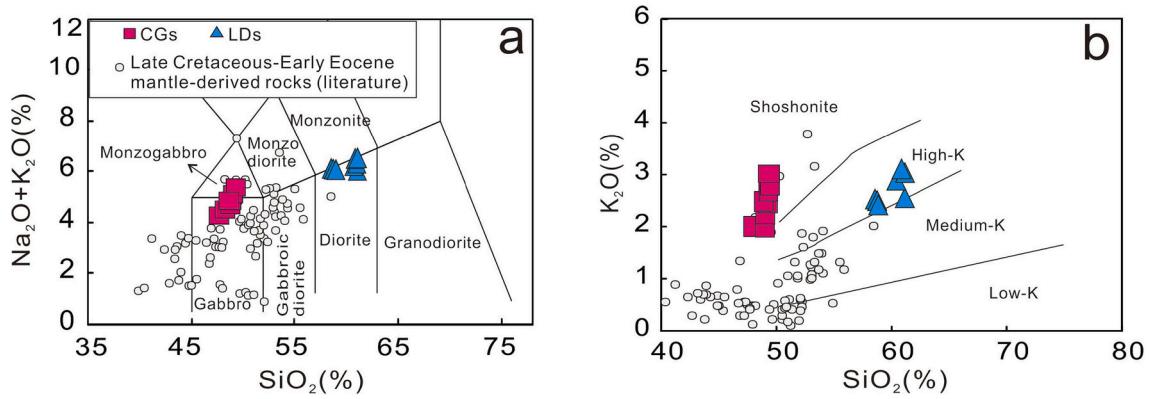


Fig. 4. (a) $\text{Na}_2\text{O} + \text{K}_2\text{O}$ (wt%) versus SiO_2 (wt%), and (b) K_2O (wt%) versus SiO_2 (wt%) plots for the CGs and LDs. All plotted data were recalculated to 100 wt% on a volatile-free basis. Data sources for Late Cretaceous–Early Eocene (100–45 Ma) mantle-derived rocks in the southern Tibet are from Mo et al. (2008); Lee et al. (2012); Ma et al. (2013a, 2013b), Qi et al. (2018), Huang et al. (2019, 2020); Wang et al. (2019).

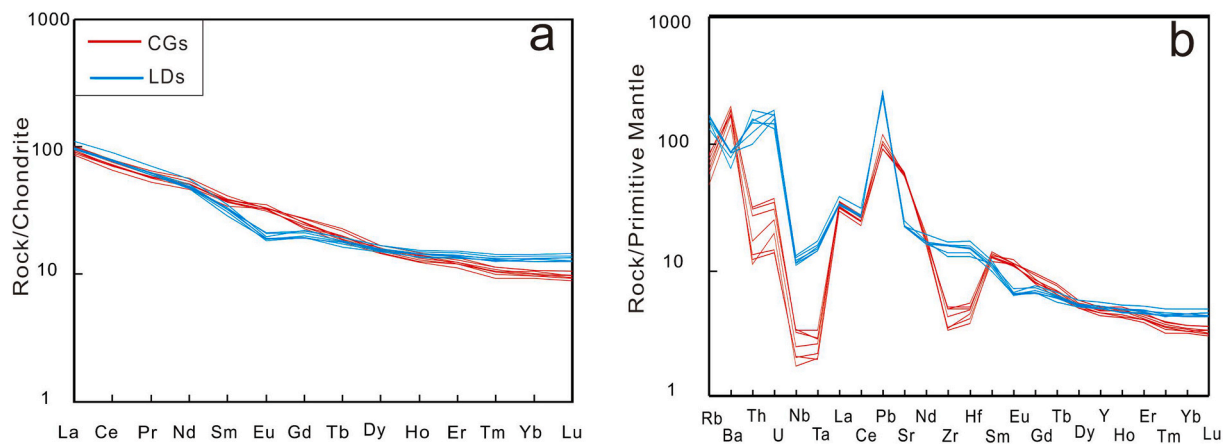


Fig. 5. (a) Chondrite-normalized REE patterns, and (b) primitive-mantle-normalized incompatible-element diagrams for the CGs and LDs. The trace-element compositions of chondrite and the primitive-mantle values are from Sun and McDonough (1989). The data sources are the same as Fig. 4.

are basaltic in composition, with low SiO_2 (48.1–49.6 wt%) and relatively high MgO (4.89–5.48 wt%) contents. They have arc-like trace element patterns with enrichment in LREEs and LILEs, and depletion in HFSEs (Fig. 5). Thus, the CGs are geochemically analogous to modern island arc basalts. There is a consensus that mafic arc magmas are

generated by melting of the depleted upper mantle enriched by materials (i.e., fluids or melts) released from the subducted slab (e.g., Elliott, 2003). The CGs have depleted Nd–Hf isotopic compositions, indicating they were produced by melting of the depleted mantle. Previous studies have proposed that the HFSEs (e.g., Nb, Ta, Zr, and Hf) and HREEs (e.g.,

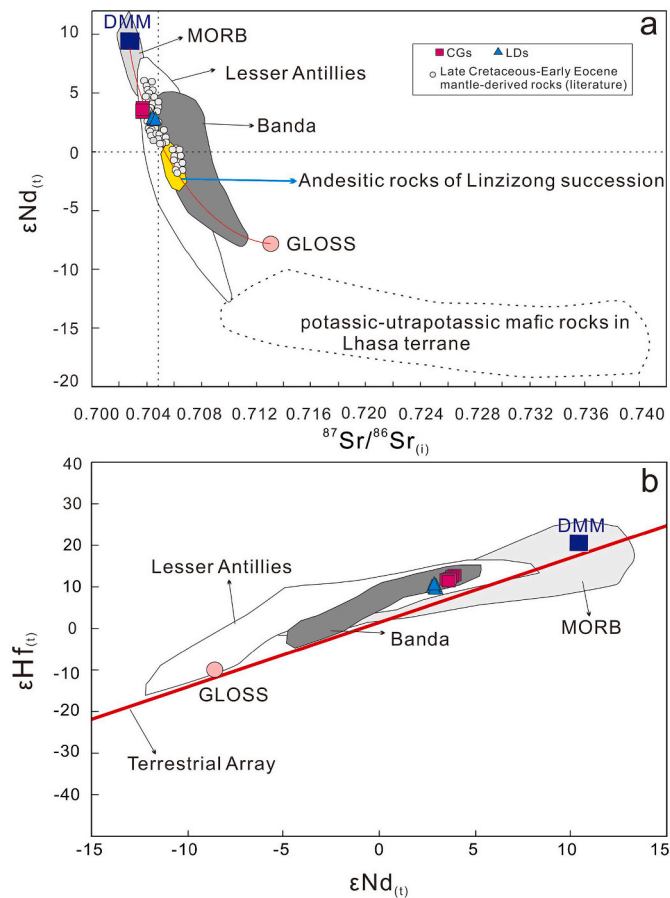


Fig. 6. Plots of (a) $\epsilon\text{Nd}_{(t)}$ versus $^{87}\text{Sr}/^{86}\text{Sr}_{(t)}$, (b) $\epsilon\text{Hf}_{(t)}$ versus $\epsilon\text{Nd}_{(t)}$ for the CGs and LDs. Data source: the fields of MORB, Less Antilles and Banda arcs are from <http://georoc.mpch-mainz.gwdg.de/georoc/>; the field of DMM (Depleted MORB Mantle) is from [Workman and Hart \(2005\)](#); the field of andesitic rocks of Linzizong succession is from [Mo et al. \(2008\)](#), and [Yan et al. \(2019\)](#); and the field of post-collisional K-rich mafic rocks in the southern Tibet is from [Turner et al. \(1996\)](#); [Miller et al. \(1999\)](#); [Ding et al. \(2003\)](#); [Gao et al. \(2007\)](#); [Guo et al. \(2013\)](#); [Liu et al. \(2015\)](#). Other data sources are the same as [Fig. 4](#).

Y and Yb) are unlikely to be transferred from the subducted slab to the mantle wedge by slab-derived fluids or melts (e.g., [Elburg et al., 2002](#); [Plank and Langmuir, 1998](#)). This is because these elements are insoluble in fluids and depleted in slab-derived melts due to the stability of rutile, other residual Ti-bearing phases (e.g., titanite), and garnet. Therefore, HFSEs and HREEs ratios (e.g., Nb/Yb, Zr/Hf, and Nb/Ta) of mafic arc magmas are likely comparable to those of the mantle wedge before modification by the slab-derived fluids and melts. The Nb/Yb (0.8–1.4) and Y/Yb (12–13) ratios of the CGs are similar to those of MORBs (Nb/Yb = 1.0–1.5; Y/Yb ~ 11; [Sun and McDonough, 1989](#); [Pfander et al., 2007](#); [Gale et al., 2013](#)). Meanwhile, Zr/Hf and Nb/Ta ratios of the CGs are also similar to those of MORBs ([Fig. 9a](#)). These features indicate that the mantle source of the CGs prior to modification by the oceanic slab (oceanic crust and oceanic sediments) was likely the depleted MORB mantle (DMM). Therefore, we assume that the elemental and Sr–Nd–Hf isotopic compositions of the pre-subduction mantle beneath southern Tibet were similar to DMM.

The mantle source of mafic arc magmas is typically enriched by fluids and/or melts released from the oceanic slab, including both the oceanic crust and overlying sediments ([Carpentier et al., 2009](#); [Labanieh et al., 2010](#)). Given that the oceanic crust is produced by melting of the depleted mantle, the fluids and/or melts released from the subducted oceanic crust usually have similar Sr–Nd–Hf isotopic compositions as MORBs ([Chauvel et al., 2009](#); [Hauff et al., 2003](#)). In contrast, oceanic

sediment-derived fluids and melts tend to have enriched Sr–Nd–Hf isotopic compositions ([Chauvel et al., 2009](#); [Plank, 2014](#); [Vervoort et al., 1999, 2011](#)). The CGs have Sr–Nd–Hf isotopic compositions that are more enriched than those of typical MORBs ([Fig. 6](#)), indicating their source contained contributions from subducted sediments rather than subducted oceanic crust ([Chauvel et al., 2009](#); [Elliott et al., 1997](#)). This is because subducted sediments are the major host for incompatible elements, and adding even a small amount of sediments will dramatically change the elemental and isotopic compositions of the mantle source ([Plank, 2014](#)). This is also consistent with the fact that the CGs plot between DMM and subducted sediments on whole-rock Sr–Nd–Hf isotope diagrams ([Fig. 6](#)), and exhibit isotopic variations similar to mafic rocks from the Lesser Antilles and Banda arcs, where the mantle wedges were strongly metasomatized by subducted sediments (e.g., [Labanieh et al., 2010](#); [Nebel et al., 2011](#)). The CGs plot within the field for global oceanic sediments on a Th/La versus Th diagram ([Fig. 9b](#)), which is also consistent with the involvement of subducted sediments in their mantle source. In addition, a subducted sediment component can account for the positive $\Delta\epsilon\text{Hf}_{(t)}$ values of these rocks. Given that zircon-free oceanic sediment-derived melts generally have more radiogenic Hf at a given Nd isotopic composition as compared with the mantle array due to the “zircon effect” (e.g., [Carpentier et al., 2009](#); [Vervoort et al., 2011](#)), mantle metasomatized by such sediment-derived melts will acquire the high $\Delta\epsilon\text{Hf}_{(t)}$ values of the sediments. In the case of sediment-derived, fluid-mediated enrichment, Nd is more fluid-mobile than Hf ([Hanyu et al., 2006](#)), and the fluids can also decouple the Hf–Nd isotopic systems ([Hanyu and Tatsumi, 2002](#); [Pearce et al., 1999](#)). Therefore, fluid-modified mantle material is likely to have less radiogenic Nd at a given Hf isotopic composition, and will also have positive $\Delta\epsilon\text{Hf}_{(t)}$ values. Thus, the sub-arc mantle metasomatized by either subducted sediment-derived melts or fluids could generate mafic rocks with positive $\Delta\epsilon\text{Hf}_{(t)}$ values, as is the case for the CGs. In summary, the mantle source of the CGs was a mixture source of DMM and recycled sediments.

5.2.2. Mantle source of the LDs

Compared with the CGs, the LDs are intermediate rocks with high SiO_2 (58.6–61.1 wt%) and low MgO (2.71–3.24 wt%) contents, and higher Mg# values (49.4–50.8) than melts directly produced by melting of the mafic crust (i.e., metabasalts; $\text{Mg}\# < 40$; [Rapp and Watson, 1995](#)). In general, intermediate rocks with high Mg# values and depleted Sr–Nd–Hf isotopic compositions can form either by the interaction between oceanic crust-derived melts and the mantle or by FC of mantle-derived mafic melts (e.g., [Defant and Drummond, 1990](#); [Grove and Christy, 2012](#); [Kelemen et al., 2003](#)). Based on two main observations, we propose that the LDs did not form by the first mechanism: (1) intermediate rocks produced in this way tend to have adakitic affinities, such as high Sr/Y ratios (>20), and low Y (<18 ppm) and Yb (<1.8 ppm) contents ([Defant and Drummond, 1990](#); [Kelemen et al., 2003](#)), in contrast to the LDs that have low Sr/Y ratios (18–22), and high Y (22.9–25.7 ppm) and Yb (2.1–2.5 ppm) contents; (2) compared with andesitic rocks of the Linzizong volcanic succession, which were produced by partial melting of Neo-Tethyan oceanic crust ([Mo et al., 2008](#)), the LDs have markedly more depleted Sr–Nd isotopic compositions ([Fig. 6a](#)). As such, the LDs most likely formed by FC of mantle-derived mafic melts. Based on the following observations, we propose that the LDs were generated by melting of a mantle source that was similar to the CGs: (1) the LDs were produced in the Gangdese continental arc setting, and exhibit arc-like trace element patterns ([Fig. 5b](#)) and Sr–Nd–Hf isotopic compositions similar to those of the CGs; (2) Zr/Hf and Nb/Ta ratios of the LDs are comparable to those of MORBs ([Fig. 9a](#)), indicating their mantle prior to modification by oceanic slab (oceanic crust and oceanic sediments)-derived fluids and melts was similar to DMM, and the high Th contents and Th/La values of the LDs require a sediment component in their mantle source ([Fig. 9b](#)). Therefore, we propose that the mantle source of the LDs was also a mixture source of DMM and recycled oceanic sediments.

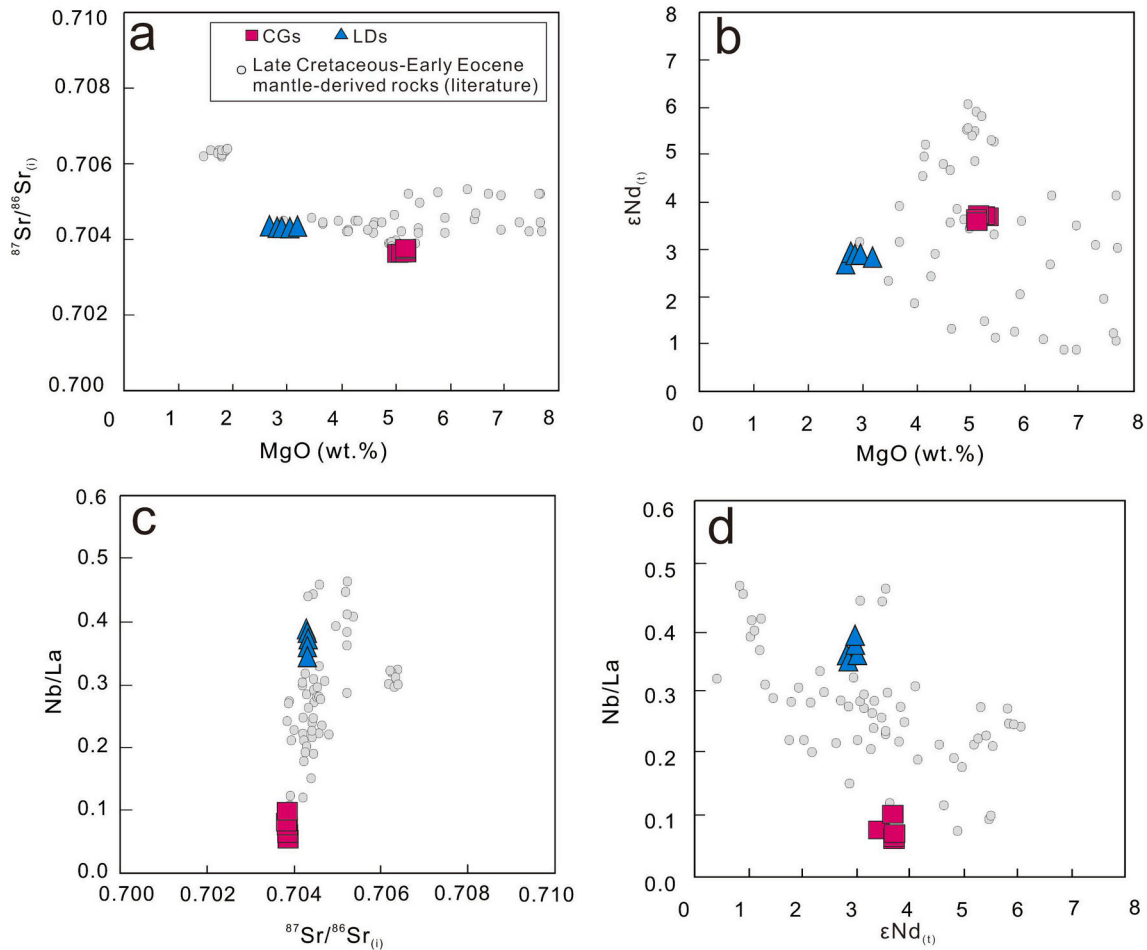


Fig. 7. Plots of (a) $^{87}\text{Sr}/^{86}\text{Sr}_{(t)}$ versus MgO (wt%), (b) $\epsilon\text{Nd}_{(t)}$ versus MgO (wt%), and (c) Nb/La versus $^{87}\text{Sr}/^{86}\text{Sr}_{(t)}$, and (d) Nb/La versus $\epsilon\text{Nd}_{(t)}$ for the CGs and LDs. The data sources are the same as Fig. 4.

5.3. Mantle enrichment by sediment-derived fluids or melts

In general, oceanic sediment-derived fluids are enriched in LILEs and Pb, whereas oceanic sediment-derived melts are characterized by significant enrichment in Th and LREEs (Elliott et al., 1997; Johnson and Plank, 1999; Pearce et al., 1999). The high Ba, U, Pb, and Sr, and low Th contents of the CGs (Fig. 5b), suggest their mantle source contains subducted oceanic sediment-derived fluids rather than melts (Johnson and Plank, 1999; Pearce and Stern, 2006). This is consistent with their high Sr/Th, Ba/La, and Ba/Th, but low Th/Nd, Th/Yb, Th/Sm, and Th/Ce ratios (Fig. 10; Woodhead et al., 1998; Hanyu et al., 2006). In contrast, the LDs are enriched in Th (Fig. 5b), denoting the involvement of subducted oceanic sediment-derived melts in their mantle source, because Th is immobile in fluids (Hermann and Spandler, 2008). The LDs have high Th/Nd, Th/Yb, Th/Sm, and Th/Ce ratios, and low Sr/Th, Ba/La, and Ba/Th ratios (Fig. 10), indicating that their mantle source was enriched by the subducted sediment-derived melts. Thus, the CGs likely represent partial melts of DMM that had been enriched by fluids from subducted sediments, whereas the LDs were produced by melting of DMM that had been enriched by melts from subducted sediments. To further estimate the proportion of subducted sediment-derived components in the mantle source of the CGs and LDs, we carried out modeling of Ba/La and Nd-Hf isotope ratios using an approach similar to Guo et al. (2016). The results indicate that the CGs could have been generated by melting of DMM enriched by 3%–4% sediment-derived fluids and the LDs could result from melting of DMM enriched by 1.5%–2.5% sediment-derived melts (Fig. 11).

Given that the CGs and LDs were sampled in the central–southern Lhasa Terrane (87–92.5°E) (Fig. 1), and in order to better constrain how the subducted slab-derived components geochemically modified the local mantle beneath southern Tibet during the late stage (100–45 Ma) of Neo-Tethyan slab subduction, we compiled previously published data for mantle-derived mafic–intermediate rocks during this time in the central–southern Lhasa Terrane (Huang et al., 2019, 2020; Lee et al., 2012; Ma et al., 2013a, 2013b; Mo et al., 2008; Qi et al., 2018; Wang et al., 2019). The Sr–Nd isotope data for these mafic–intermediate rocks plot along a mixing line between DMM and subducted sediments, and have a large range of Sr–Nd isotope ratios similar to mafic igneous rocks from the Lesser Antilles and Banda arcs (Fig. 6). This indicates that subducted oceanic sediments were present in their mantle source. Meanwhile, the Sr/Th, Ba/La, Ba/Th, Th/Ce, and Th/Sm ratios of these mafic–intermediate rocks plot between those of the CGs and LDs (Fig. 10), also suggesting the involvement of subducted sediments in their source was in the form of fluids and/or melts. Therefore, intensive dehydration and melting of subducted oceanic sediments occurred in the late stage (100–45 Ma) of Neo-Tethyan slab subduction beneath the study area.

5.4. Thermal events responsible for generating the CGs and LDs

The Lhasa Terrane records the subduction of the Neo-Tethyan oceanic lithosphere, subsequent collision of the Indian and Asian continents, and subduction of Indian continental lithosphere (e.g., Chung et al., 2005). These events were accompanied by a series of geodynamic

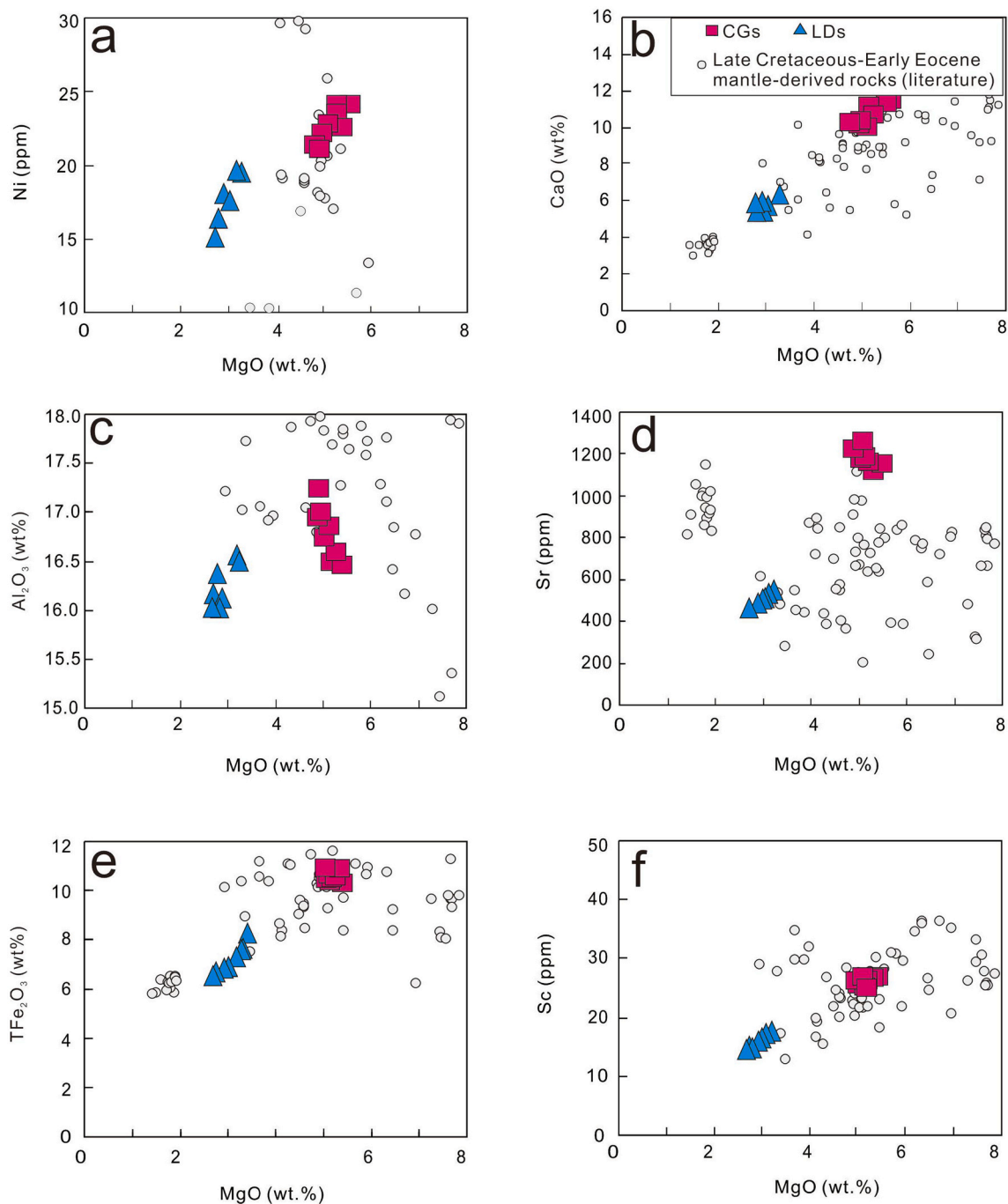


Fig. 8. Variations in (a) Ni (ppm) versus MgO (wt%), (b) CaO (wt%) versus MgO (wt%), (c) Al₂O₃ (wt%) versus MgO (wt%), (d) Sr (ppm) versus MgO (wt%), (e) TFe₂O₃ (wt%) versus MgO (wt%), and (f) Sc (ppm) versus MgO (wt%) for the CGs and LDs, respectively. The data sources are the same as Fig. 4.

processes, including oceanic ridge subduction, slab rollback, and break-off of the subducted oceanic plate (e.g., Chung et al., 2005; Zhu et al., 2013). Based on the occurrence of high-temperature charnockites and granulite-facies metamorphism of the surrounding metamorphic rocks at 90 ± 5 Ma, it has been suggested that Neo-Tethyan oceanic ridge subduction may have occurred during 100–90 Ma (e.g., Zhang et al., 2010). The abrupt decrease in the convergence rate between India and Asia, exhumation of ultra-high-pressure rocks, exposure of ocean island basalt (OIB)-like mafic rocks, and rapid addition of juvenile material in the Lhasa Terrane during 53–45 Ma suggest that slab break-off of the subducted Neo-Tethyan slab may have taken place during this period (e.g., Ji et al., 2016; Leech et al., 2005; Zhu et al., 2013, 2015). Both ridge subduction and slab break-off can rapidly alter the thermal state of

a subducted slab due to the upwelling of hot asthenosphere (Davies and von Blanckenburg, 1995; Thorkelson and Breitsprecher, 2005). The ascending hot asthenosphere can heat an overlying subduction zone and result in extensive melting and dehydration of oceanic sediments. Therefore, dehydration and melting of subducted oceanic sediments beneath the study area may have been ubiquitous in the late stage (100–45 Ma) of Neo-Tethyan oceanic slab subduction, which is a common phenomenon in subduction systems with a high geothermal gradient (e.g., Watt et al., 2013).

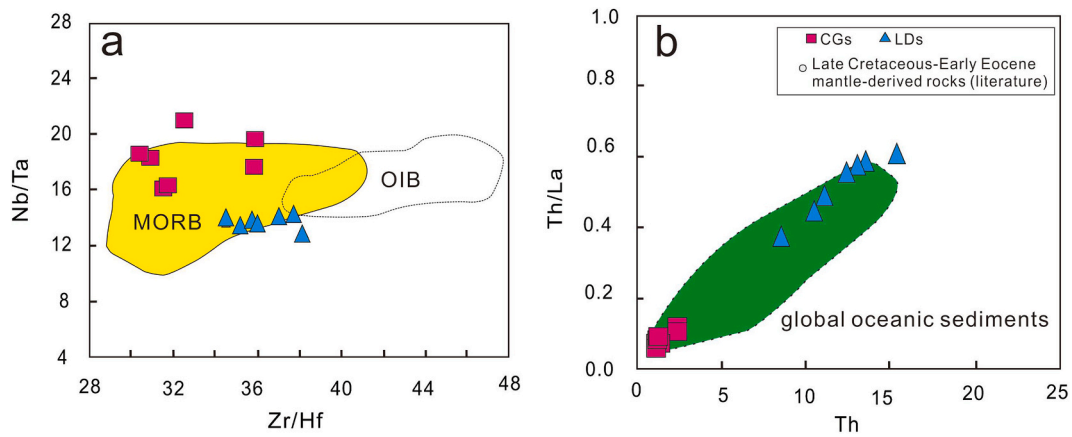


Fig. 9. Diagrams of (a) Nb/Ta versus Zr/Hf, and (b) Th/La versus Th (ppm) for the CGs and LDs. Data sources are as follows: the field of MORB are from Sun and McDonough (1989), Workman and Hart (2005); and Pfander et al. (2007); the field of OIB is from Sun and McDonough (1989); Pfander et al. (2007); and Gale et al. (2013) the field of global sediments is from Plank (2014).

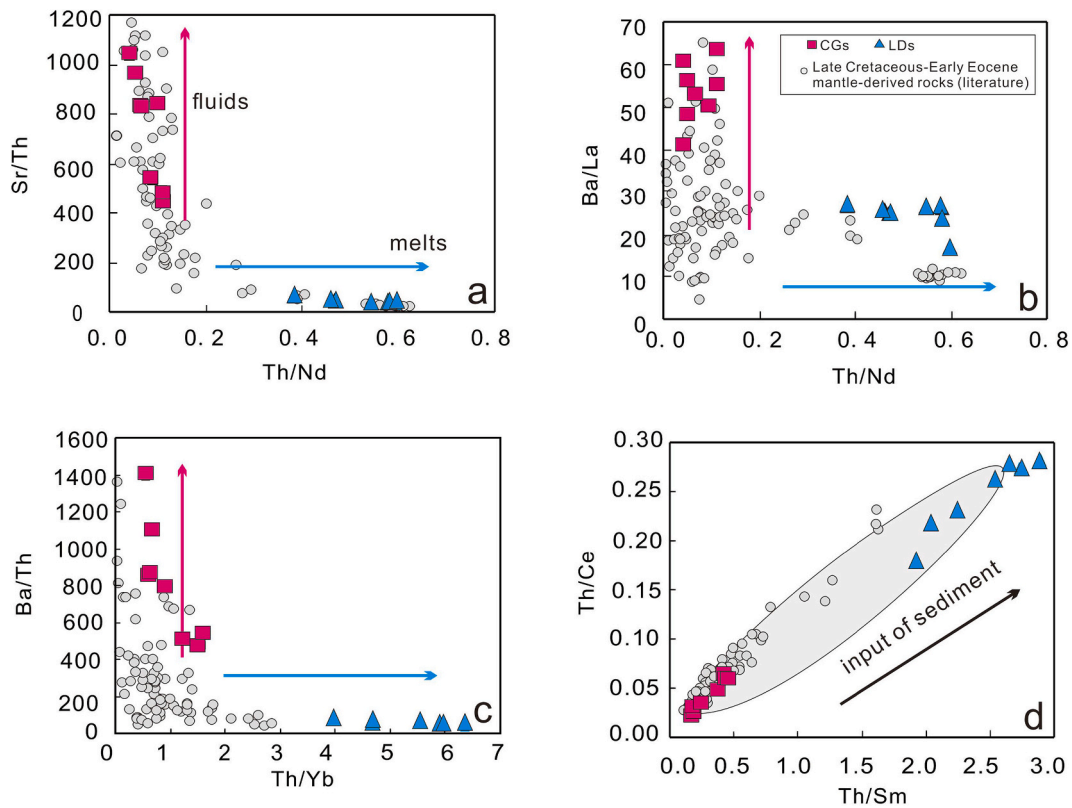


Fig. 10. Plots of (a) Sr/Th versus Th/Nd, (b) Ba/La versus Th/Nd, (c) Ba/Th versus Th/Yb, and (d) Th/Ce versus Th/Sm for the CGs and LDs. The data sources are the same as Fig. 4.

5.5. Implications for the Miocene isotopically enriched mantle beneath the Lhasa Terrane

The Miocene post-collisional K-rich mafic rocks are widely distributed in the western Lhasa Terrane (west of 87°E) and have highly enriched Sr-Nd isotopic compositions, indicating they formed by melting of an isotopically enriched mantle source. This isotopically enriched mantle source is argued to have been generated by the subduction of either the Neo-Tethyan oceanic slab or Indian continental slab (e.g., Ding et al., 2003; Gao et al., 2007; Tommasini et al., 2010; Guo et al., 2013; Liu et al., 2015).

As we discussed in Section 5.2, the addition of sediments can

dramatically change the elemental and isotopic compositions of mantle source. Our study has shown that melting and/or dehydration of oceanic sediments occurred mainly during the late stage (100–45 Ma) of Neo-Tethyan oceanic slab subduction. This means that the mantle beneath the Lhasa Terrane was most likely enriched during 100–45 Ma. The CGs and LDs, and previously studied mafic-intermediate igneous rocks during this time in the central-southern (87–92.5°E) Lhasa Terrane, have depleted Sr-Nd isotopic compositions, which are distinct from the Miocene post-collisional K-rich mafic rocks that have enriched Sr-Nd isotopic compositions (Fig. 6). This indicates that the mantle beneath the Lhasa Terrane (at least the central-southern Lhasa Terrane) was still depleted, even though the mantle source had been enriched during

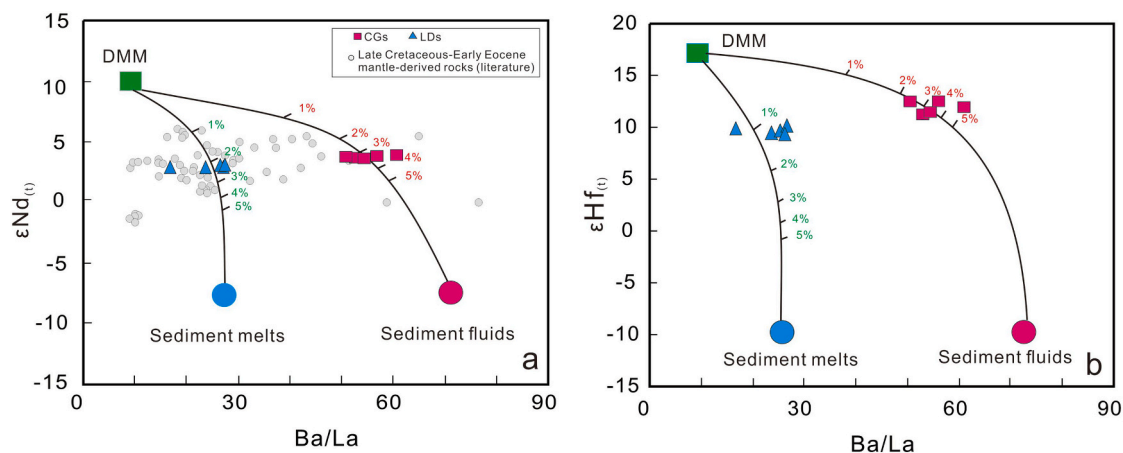


Fig. 11. Diagrams of (a) $\epsilon\text{Nd}_{(t)}$ versus Ba/La, and (b) $\epsilon\text{Hf}_{(t)}$ versus Ba/La for CGs and LDs. Date source: the data of DMM (Ba = 3, La = 0.4, $\epsilon\text{Nd}_{(t)} = +18$, and $\epsilon\text{Hf}_{(t)} = +10$) are from Workman and Hart (2005). The data of the subducted sediment (Ba = 720, La = 26, $\epsilon\text{Nd}_{(t)} = -8$, and $\epsilon\text{Hf}_{(t)} = -10$) are from Plank (2014). The bulk partition coefficients (KD) during sediment melting, the elemental mobility in aqueous fluid and the dehydration volume released from sediment are followed from Guo et al. (2016). Other data sources are the same as Fig. 4.

subduction of the Neo-Tethyan oceanic slab. This demonstrates that simply melting of the mantle source that had been enriched by oceanic sediments beneath the Lhasa Terrane could not have produced the K-rich mafic rocks. To account for the enriched component of K-rich mafic rocks in the Lhasa Terrane, we invoke that there may be at least two possibilities: (1) prior to Neo-Tethyan oceanic slab subduction, the mantle beneath the central-southern and western Lhasa Terrane had distinct geochemical characteristics. For example, prior to Neo-Tethyan oceanic slab subduction, the mantle beneath the western Lhasa Terrane may have been an ancient Proterozoic isotopically enriched mantle as proposed in previous studies (e.g., Miller et al., 1999; Turner et al., 1996), in contrast to the DMM beneath the central-southern Lhasa Terrane as inferred in this study. This ancient isotopically enriched mantle beneath the western Lhasa Terrane was the enriched source of the K-rich mafic rocks; (2) given that underthrusting of Indian continental lithosphere followed the India–Asia collision (Freymueller, 2011), it is apparent that Indian continent crust may have metasomatized the source region of the K-rich mafic rocks (Ding et al., 2003). However, it is beyond the scope of this study to demonstrate whether ancient mantle or Indian continent crust was the enriched component in the source of the K-rich mafic rocks. Our study clearly stresses that the enriched component of these K-rich mafic rocks cannot be simply attributed to materials derived from the Neo-Tethyan oceanic slab, and other components are required to account for the enriched source of these K-rich mafic rocks.

6. Conclusions

Based on new geochronological, geochemical, and Sr–Nd–Hf isotope data for two plutons (CGs and LDs), along with previously published data, we examined how the mantle was geochemically modified during the late stage (100–45 Ma) of Neo-Tethyan oceanic slab subduction beneath the Lhasa Terrane. The geochemical characteristics of the CGs and LDs indicate that they were produced by melting of DMM that had been enriched by oceanic sediment-derived fluids and melts, respectively. Our and previous studies suggest that dehydration and melting of subducted Neo-Tethyan oceanic sediments may have commonly occurred in the late stage (100–45 Ma) of Neo-Tethyan oceanic slab subduction beneath the study area. This is a common phenomenon in relatively hot subduction zones, such as the Lesser Antilles and Banda arcs. Although oceanic sediments would have dramatically changed the elemental and isotopic compositions of the mantle source, the mantle beneath the Lhasa Terrane was still isotopically relatively depleted as compared with the mantle source of post-collisional K-rich mafic rocks

in the Lhasa Terrane. This demonstrates that the enriched component in the mantle source of the K-rich mafic rocks cannot be simply attributed to Neo-Tethyan oceanic slab subduction, and requires additional materials in their mantle source.

Declaration of Competing Interest

The authors declare that they have no competing financial interests or personal relationships that could have appeared to influence the work in this paper.

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Appendix A. Supplementary data

Supplementary data to this article can be found online at <https://doi.org/10.1016/j.lithos.2021.106505>.

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